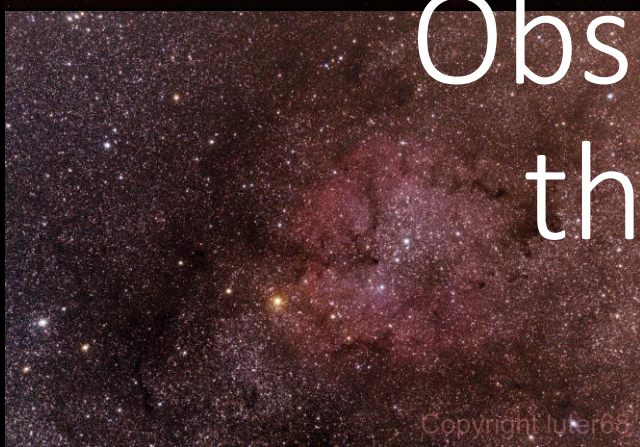
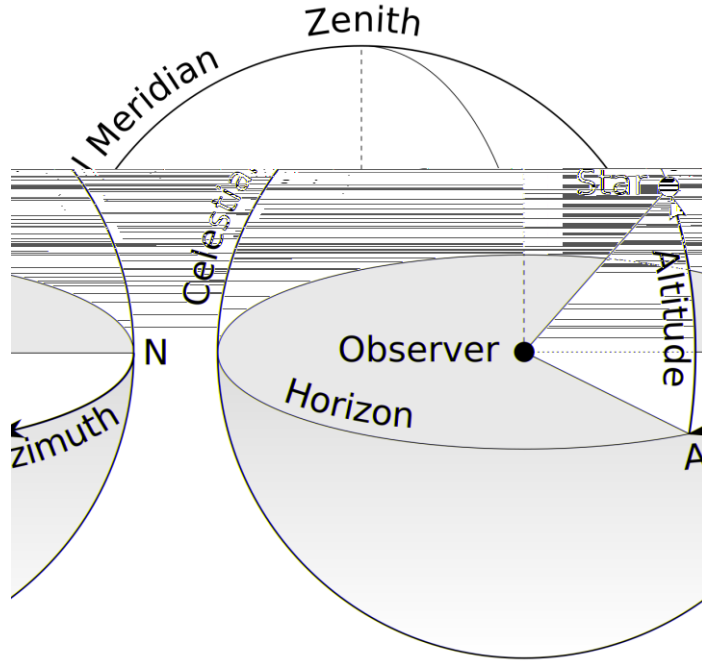


Observing and Photographing the Night Sky – First Steps

L. Teriaca

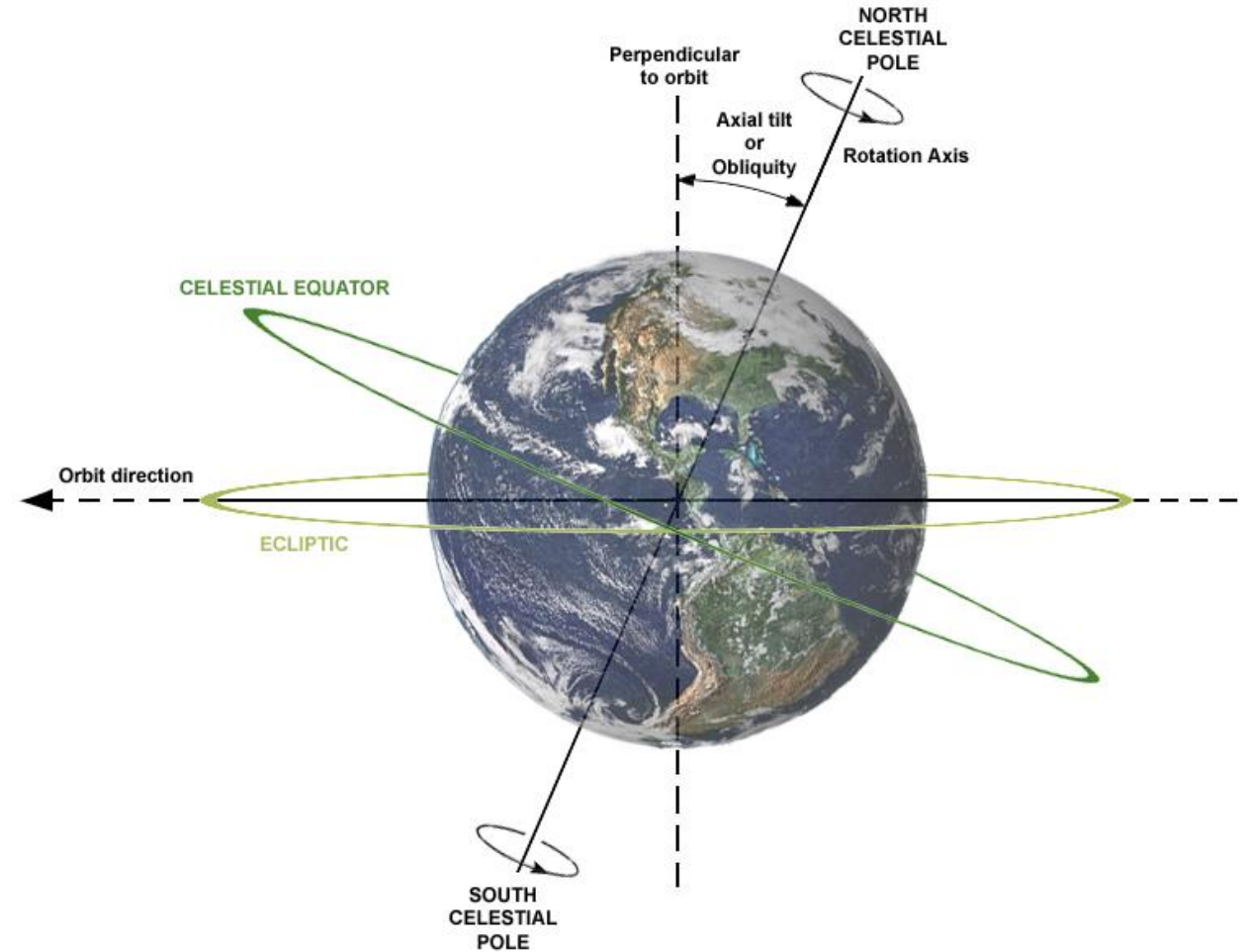


Coordinate Systems



Alt-Azimuthal

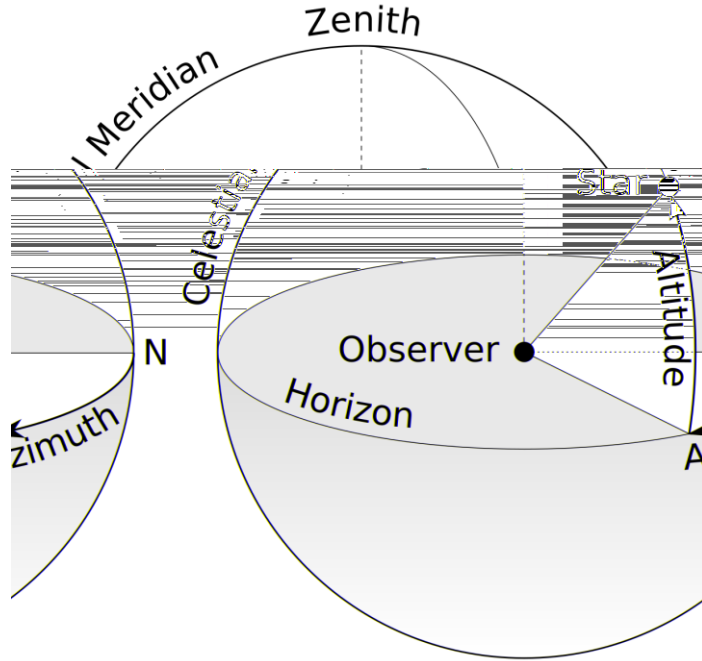
- Intuitive
- Depends on the observer's location



Equatorial

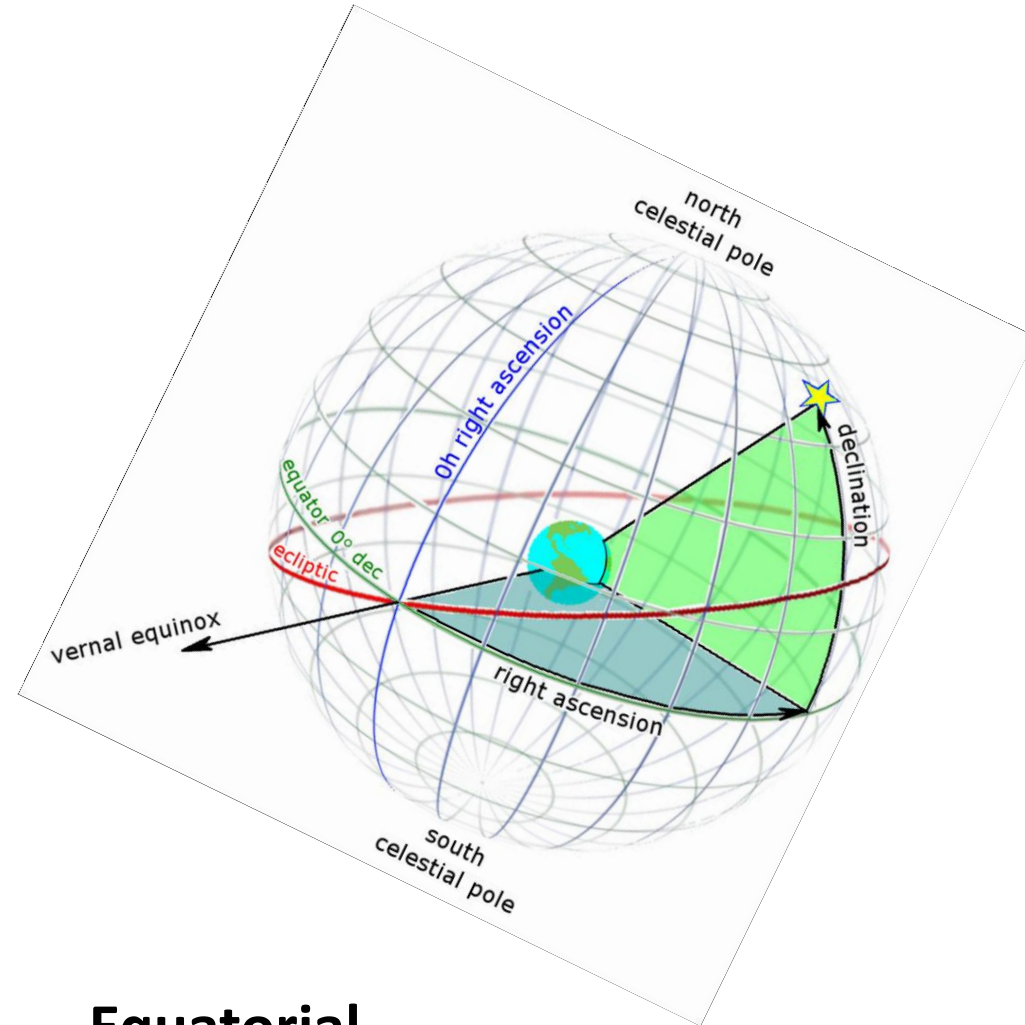
- Independent from observer's location

Coordinate Systems



Alt-Azimuthal

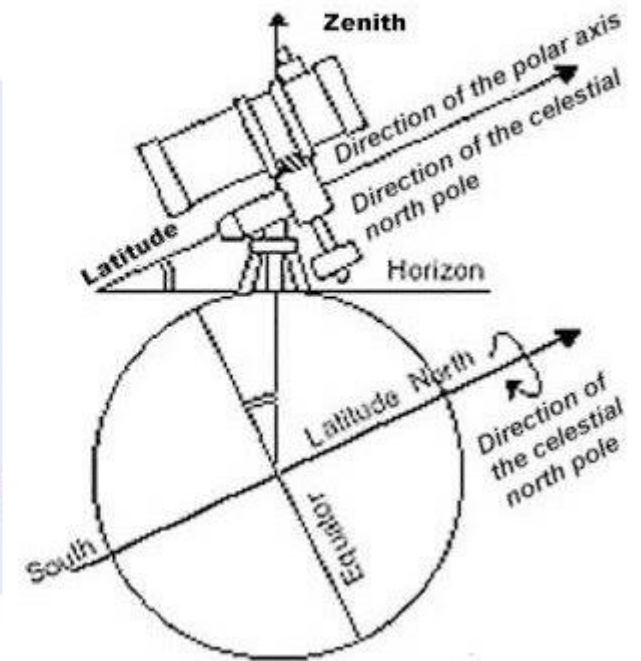
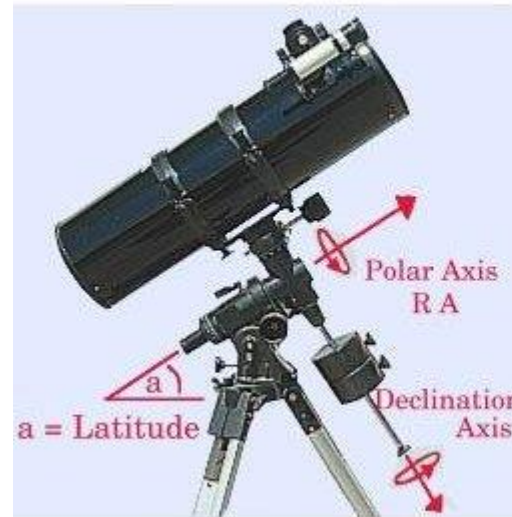
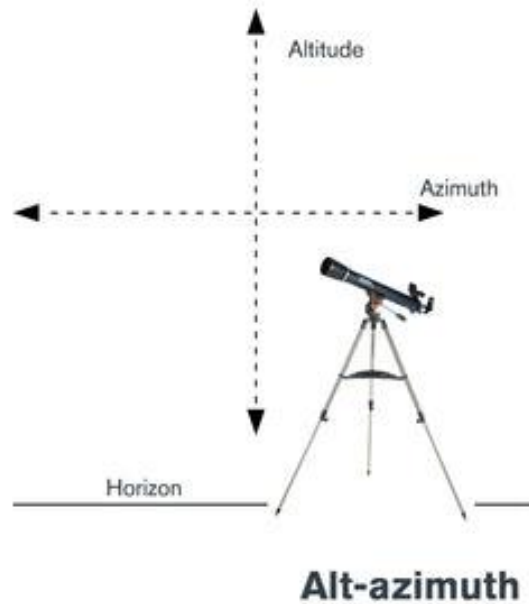
- Intuitive
- Depends on the observer's location



Equatorial

- Independent from observer's location
- It slowly changes over time due to precession

Mount types



Alt-Azimuthal

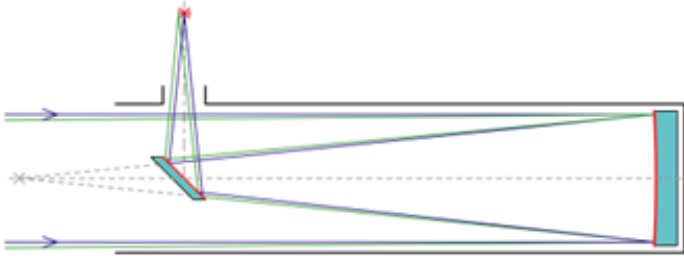
- Simple and fast to set up
- Good for visual and Skyscapes
- OK for Planetary High Res.
 - Good when goto.

Equatorial (German Eq. Mount or GEM shown. There are others!)

- Complex to set up (requires polar alignment)
- Necessary (generally) for long exposure astrophotography

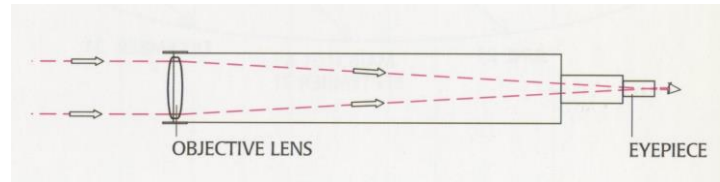
Main telescope types

Reflectors (Newtonians)



- Use only mirrors
- Newtonians (simplest) can be very cost effective
- Can be awkward for visual on a GEM
- Become quickly very large and heavy
- Require a coma corrector when fast

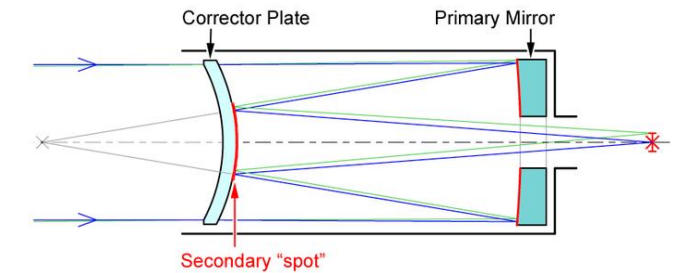
Refractors



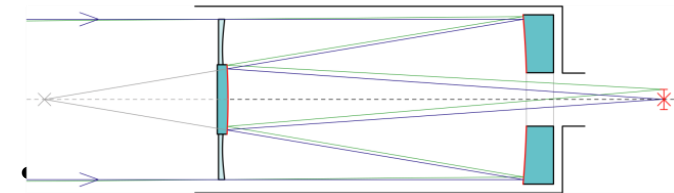
- Use only lenses
- Quite versatile
- Large diameters ($>5''$) prohibitive
- Excellent for astrophotography (higher quality models)
- Require diagonal for visual
- Require flattener for imaging

Catadioptric

Maksutov-Cassegrain



Schmidt-Cassegrain



- Generally not good for wide field
- Very compact and easier to mount (by equal diameter)
- Excellent for planetary imaging
- Require diagonal for visual

A few useful formulas and expressions (1/2)

Telescopes (and camera objectives, where you, practically, can change D)

- D = diameter of the objective
- F = focal length of the telescope
- F/D = “speed” of the telescope. Also indicated as $F/\#$ (e.g., $D=100$ mm, $F=500$ mm is $F/D=5$ or $F/5$)
 - The smaller, the brighter the image, the most demanding the manufacturing
- Diffraction limited angular resolution $\theta = 1.22 \lambda / D$

Binoculars

- Characterized by magnification x diameter of lenses ($M \times D$). E.g., 8x42 has 42 mm lenses and provides 8x magnification
- D/M = exit pupil.
 - Human pupil at night is 7 mm max (in kids).
 - Older people get smaller pupils at night (≈ 5).
 - Exit pupils between 4 and 6 are good.
- Be aware of diameter fever! Things get heavy real fast!
- Other things are important too. Ask if interested!



20x70 (3.5)



A few useful formulas and expressions (2/2)

Telescopes – visual use

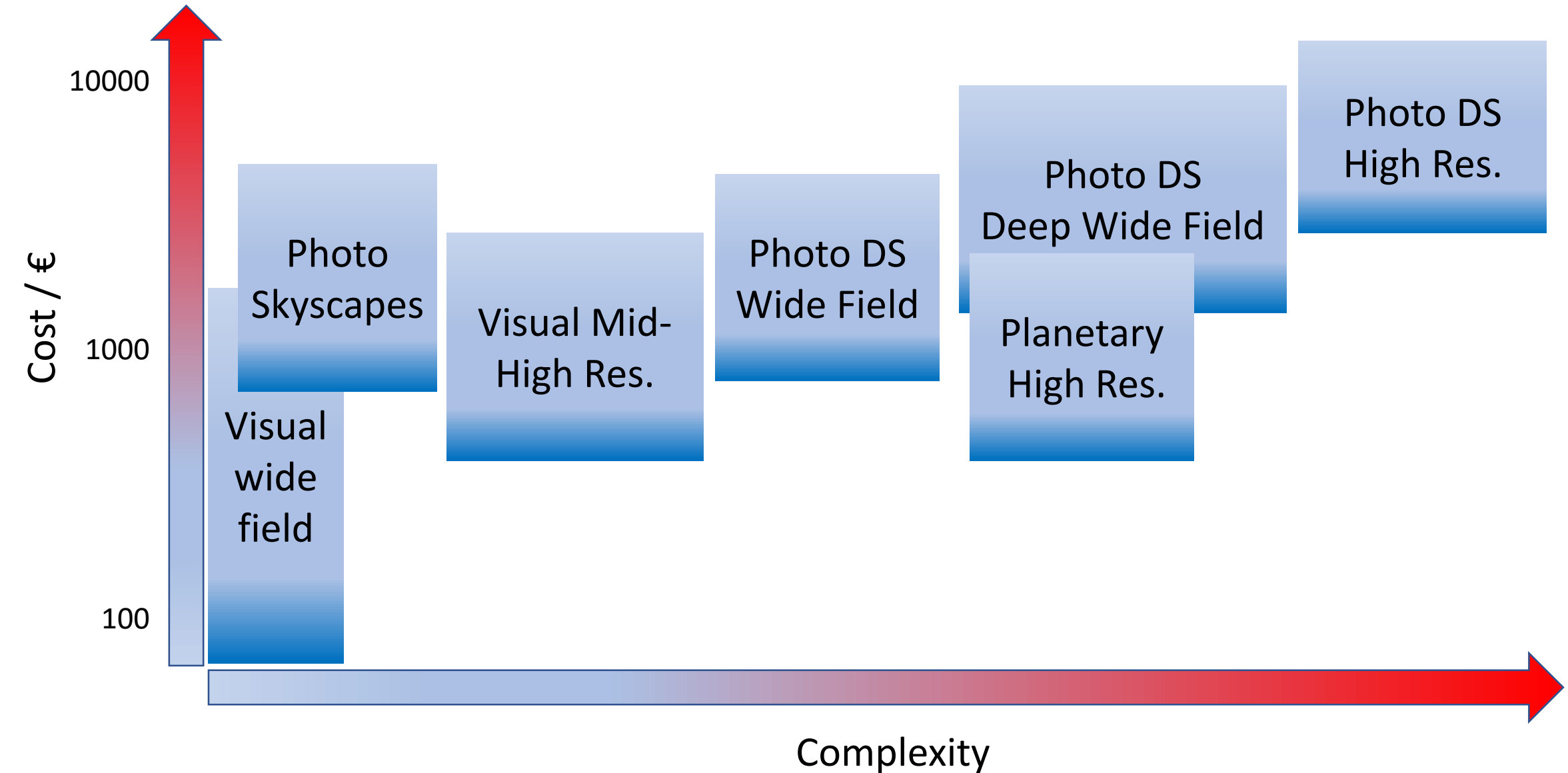
- Eyepieces are characterized by their focal length f in mm (always on barrel) and by the apparent FOV (AFOV) in degrees
- $M = F/f$ is the magnification. A 9 mm eyepiece on a telescope with $F=900$ mm magnifies 100 times.
 - Maximum magnification usually useable is $M = D$ (in mm)
 - Occasionally (good seeing) and with quality optics M can go up to $2D$ or more
- The True FOV (TFOV) can be estimated as $TFOV = AFOV / M$ (however, real TFOV can be smaller)
 - You want at least one eyepiece with a large TFOV (large AFOV to start, low M = large f)
- Exit pupil = D / M . Good values for weak extended objects between 3 and 5
- Other things are important too. Ask if interested!

Telescope - imaging

- Signal from resolved objects $N \text{ (phot s}^{-1}) \approx L \text{ (phot m}^{-2} \text{ s}^{-1} \text{ sr}^{-1}) D^2 p^2 / F^2$, where p is the pixel size
 - For a given object, if p is fixed, $N \text{ (phot s}^{-1}) \approx (D / F)^2$
 - For a given object, if you can reach diffraction limit ($p = 1.22 \lambda F / D$) as in space, $N = \text{const.}$
 - For a given object, in Earth's atmosphere (typical seeing s between 1" and 2"), $p = s F$ and $N \approx D^2$
 - Of course, extended objects do not have uniform radiance, resolution (hence contrast) increases with D until seeing kicks in

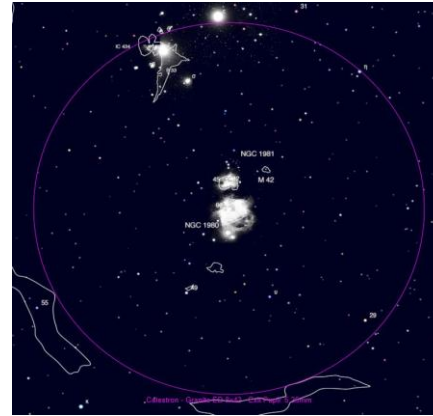
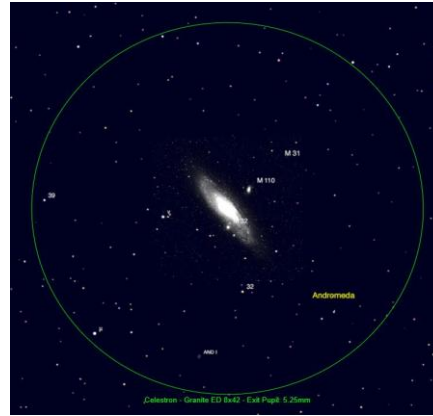
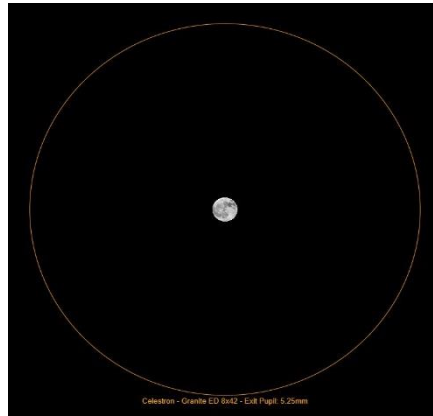


Cost and Complexity

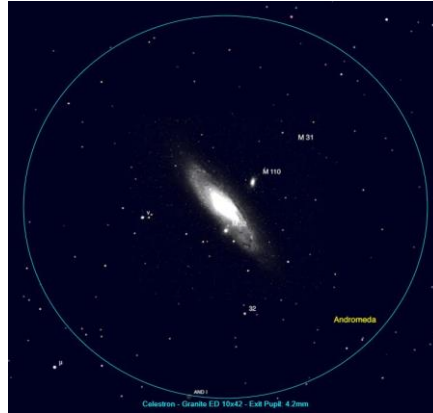
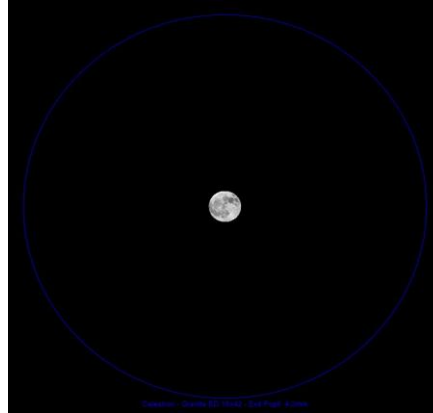


Visual Wide Field

8x42
8.1°



10x42
6.5°

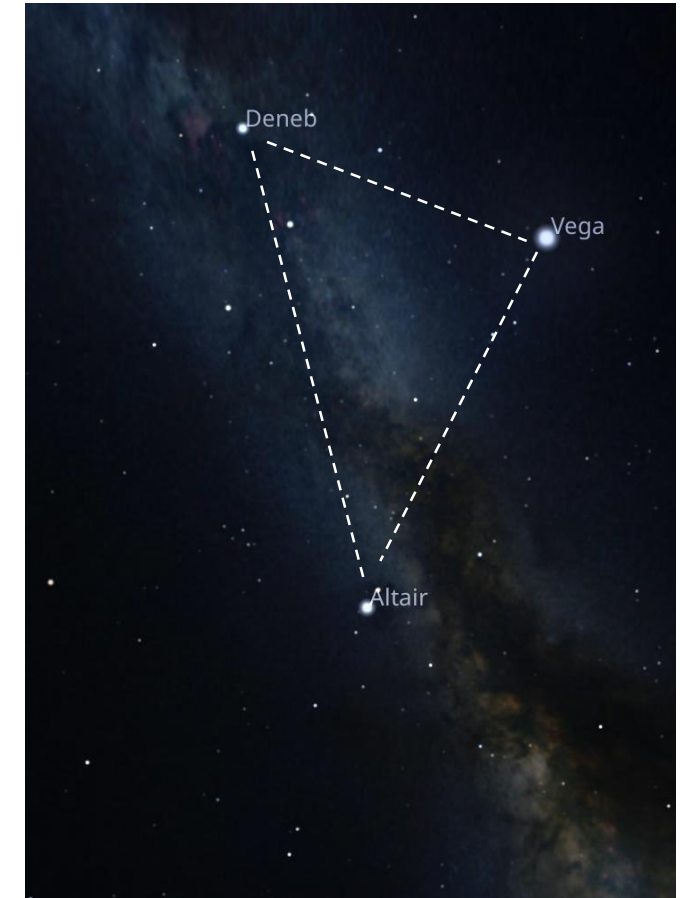


- Definitively the less demanding – naked eye observing is for free!
- Using (small) binocular can be done without a tripod
 - An 8x42 or a 10x42 are very good
- Short F refractors on an Alt-Az mount can also be a (more expensive option)

Visual Wide Field



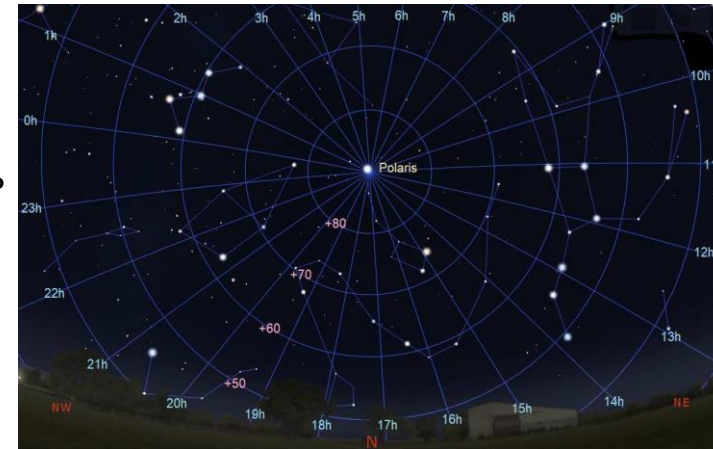
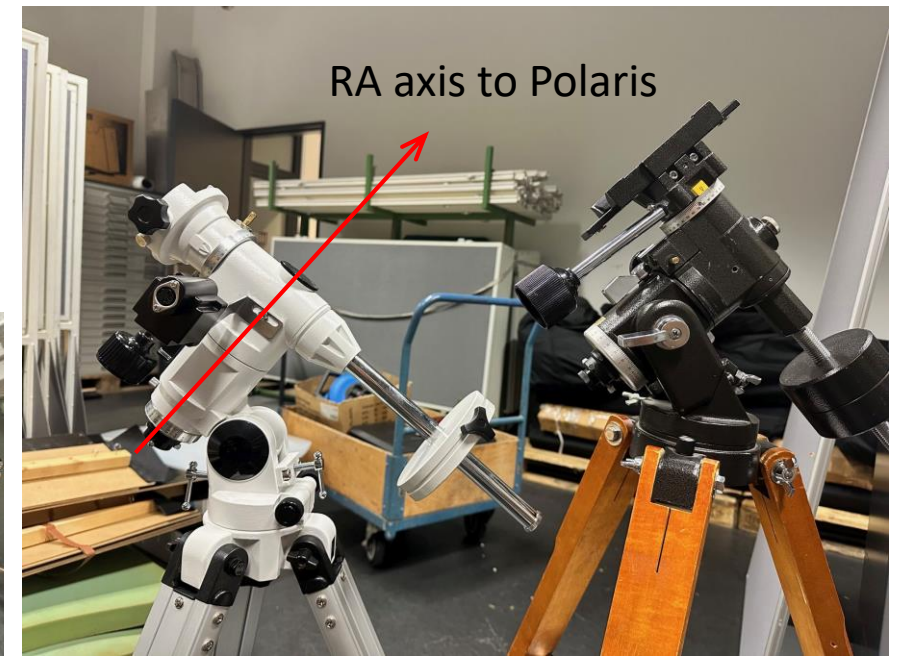
M31: Andromeda



- Sky on September 4th at 22:00.
- What you can see with naked eyes?: Milky way, Saturn, Summer triangle, constellations, (Andromeda, Mizar-Alcor)
- What with a binocular? Milky Way star fields, Andromeda, M13, Mizar-Alcor

Visual Mid-High Res.

- MPS Newton 114/900 on GEM (but generally valid for any scope on a GEM)
- Align finder to main scope (easier in daylight)
- Level tripod (put bubble level on tray)
 - doesn't need to be very precise, just not too bad
- Set local latitude using elevation knobs
- Polar align: it is the most important step
 - Point R.A. axis towards Polaris (use the azimuth knobs)
 - If you now point scope to Polaris, Dec should read $+90^\circ$
- Point a bright star not too far from the target
 - Unlock the AR circle (lower numbers for north hemisphere) and set it to the star AR. Dec should be roughly right. Lock the AR circle.
- Now move the scope to your target by matching AR & Dec values

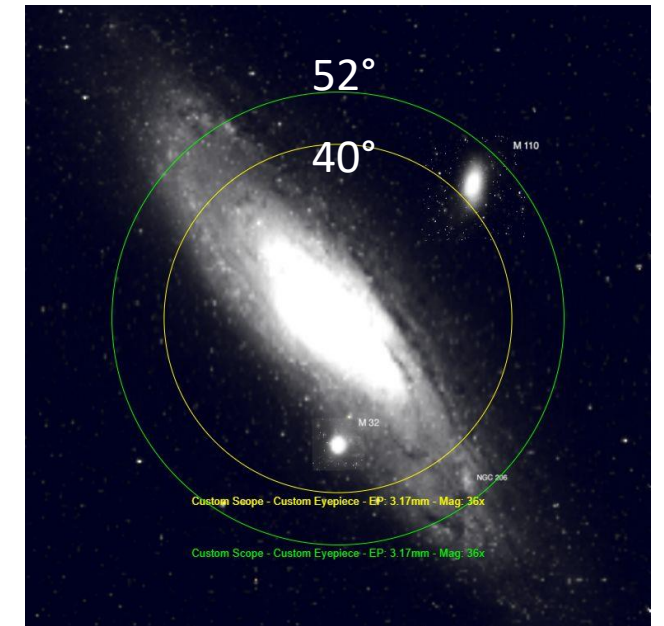


Visual Mid-High Res.

- MPS Newton 114/900 on GEM - Eyepieces

Telescope D = 114 mm, F = 900 mm			
f	M	Pupil	TFOV
25	36	3.16	1.1
12.5	72	1.58	0.55
9	100	1.14	0.4
4	223	0.51	0.18

- I do not know the AFOV of these eyepieces.
 - These are orthoscopic ones, so it is about 40° (not much)
- Use the finder first, then the 25mm eyepiece (see figure)
- Orthos have little eye relief. It may be a problem for people with glasses...





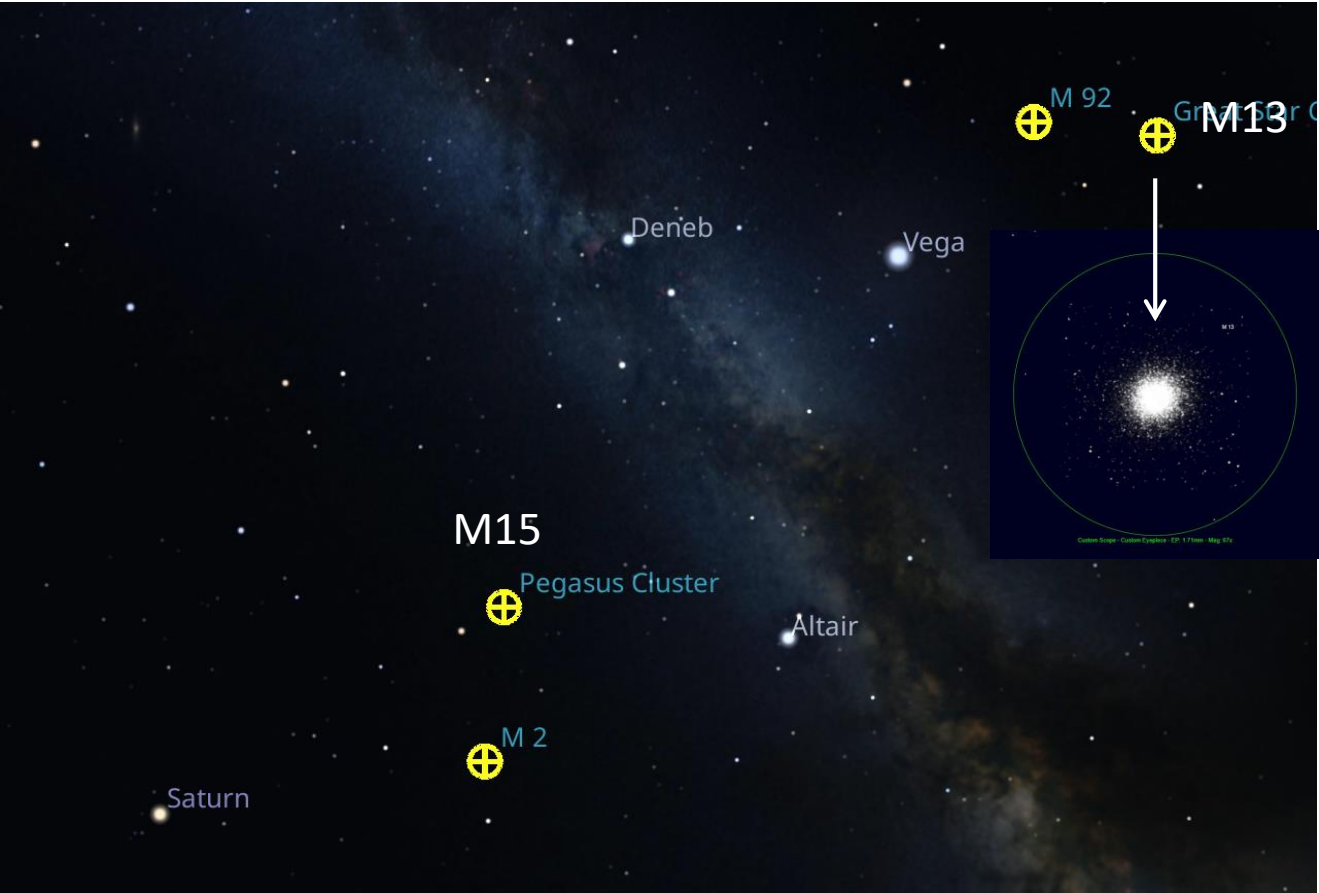
Globular clusters

Visual Mid-High Res.

Cluster	AR	Dec
M2	21 ^h 33 ^m 27.02 ^s	-00° 49' 23.7"
M13	16 ^h 41 ^m 41.24 ^s	+36° 27' 35.5"
M15	21 ^h 29 ^m 58.33 ^s	+12° 10' 01.2"
M92	17 ^h 17 ^m 07.39 ^s	+43° 08' 09.4"

Use first the 25mm to find the clusters, then try the 12.5 or the 9mm (see insert)

Cluster	AR	Dec
Altair	19 ^h 50 ^m 47.00 ^s	+08° 52' 05.96"
Vega	18 ^h 36 ^m 56.34 ^s	+38° 47' 01.28"
Deneb	20 ^h 41 ^m 25.9 ^s	+45° 16' 49"



- What with the 114/900? **M13, M15, M2, M92**, M31, M81, M27, M57, h&x Persei, Mizar, 61 Cygni, (M33, M51, M101)

Spiral galaxies



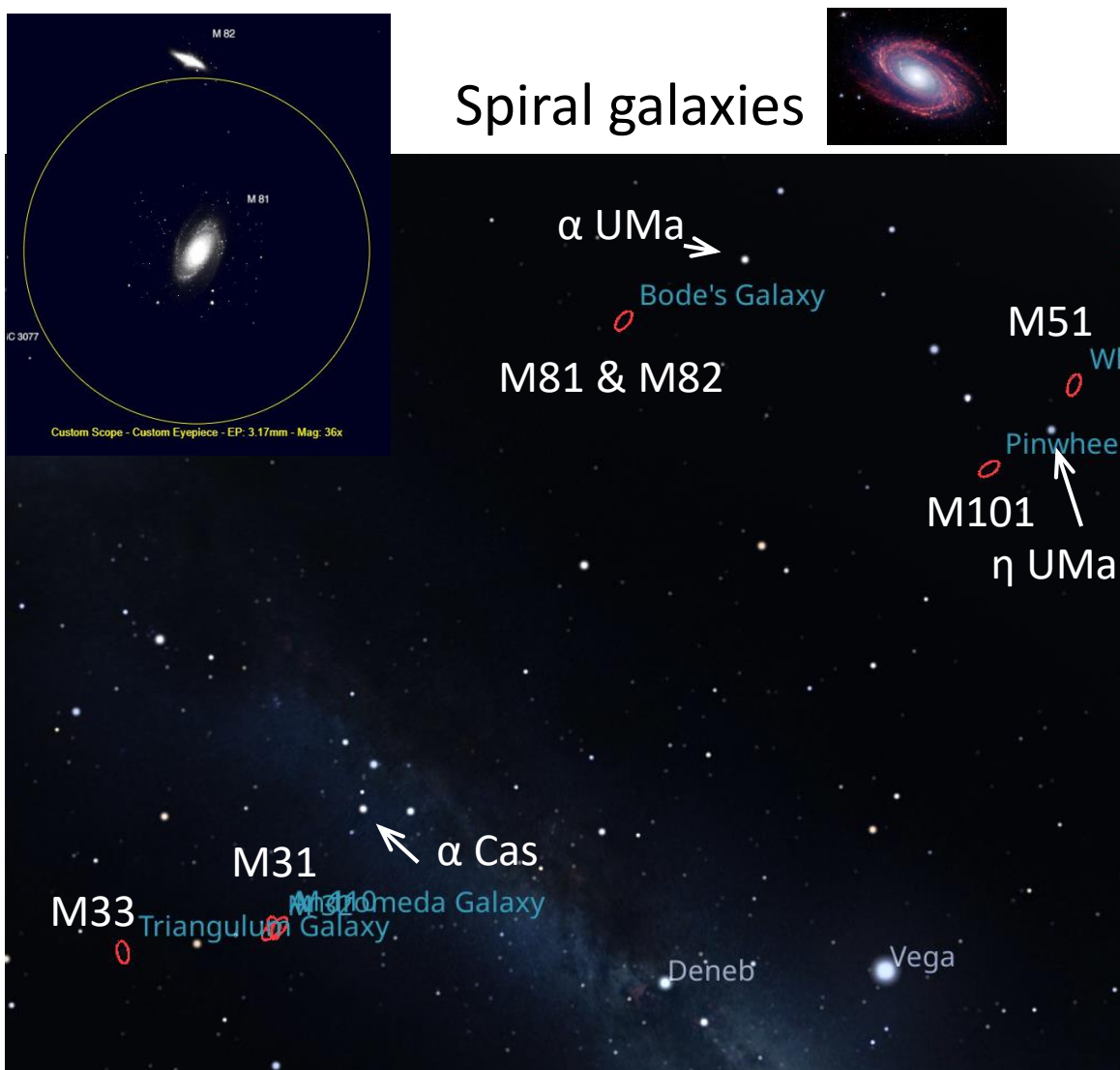
Visual Mid-High Res.

Galaxy	AR	Dec
M31	00 ^h 42 ^m 44.3 ^s	+41° 16' 09"
M33	01 ^h 33 ^m 50.02 ^s	+30° 39' 36.7"
M51	13 ^h 29 ^m 52.7 ^s	+47° 11' 43"
M81	09 ^h 55 ^m 33.2 ^s	+69° 03' 55"
M101	14 ^h 03 ^m 12.6 ^s	+54° 20' 57"

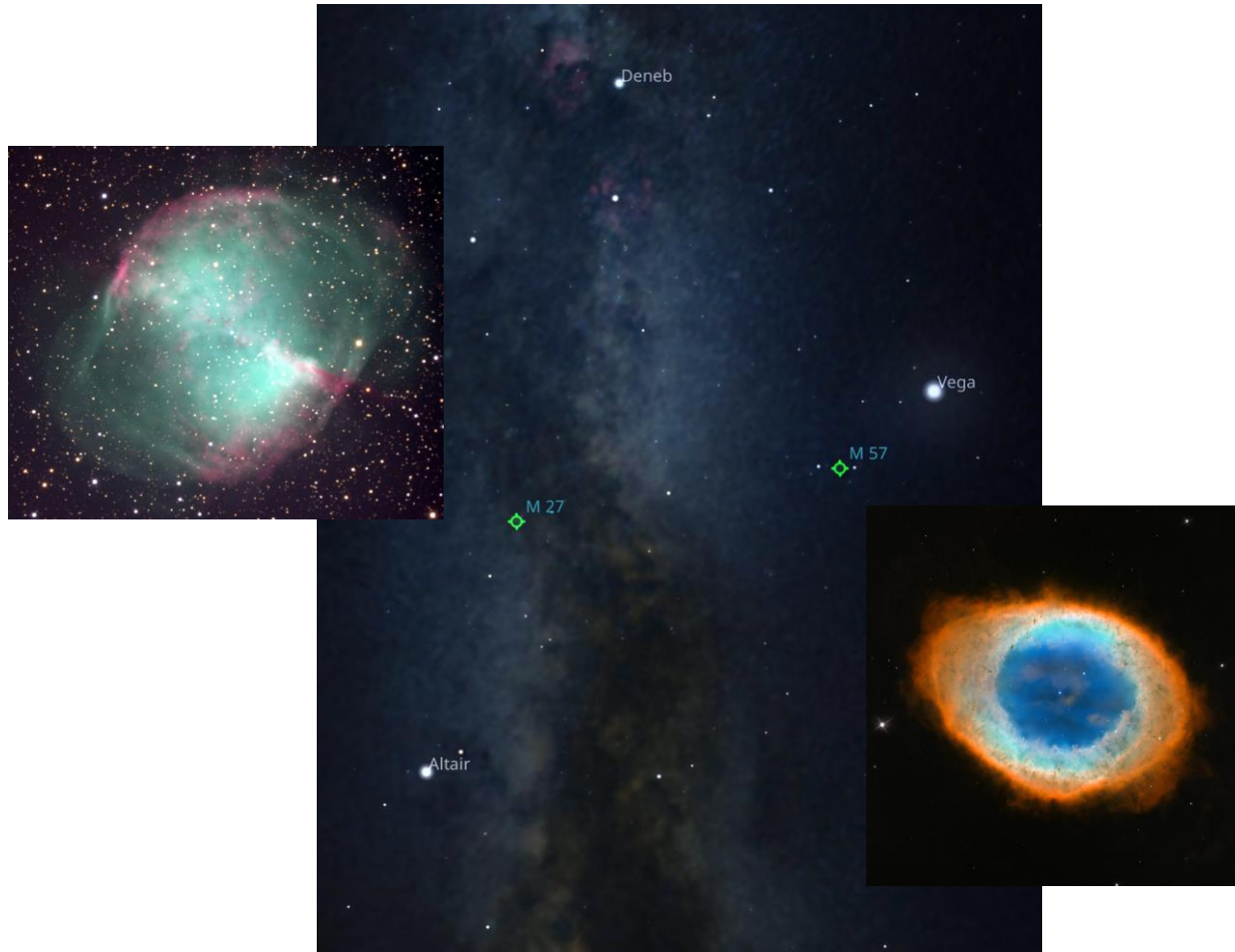
Just use the 25mm

Cluster	AR	Dec
α UMa	11 ^h 03 ^m 43.67 ^s	+61° 45' 03.72"
η UMa	13 ^h 47 ^m 32.44 ^s	+49° 18' 47.76"
α Cas	00 ^h 40 ^m 30.44 ^s	+56° 32' 14.39"

- What with the 114/900? M13, M15, M2, M92, **M31**, **M81**, M27, M57, h&x Persei, Mizar, 61 Cygni, (**M33**, **M51**, **M101**)



Planetary Nebulae



Visual Mid-High Res.

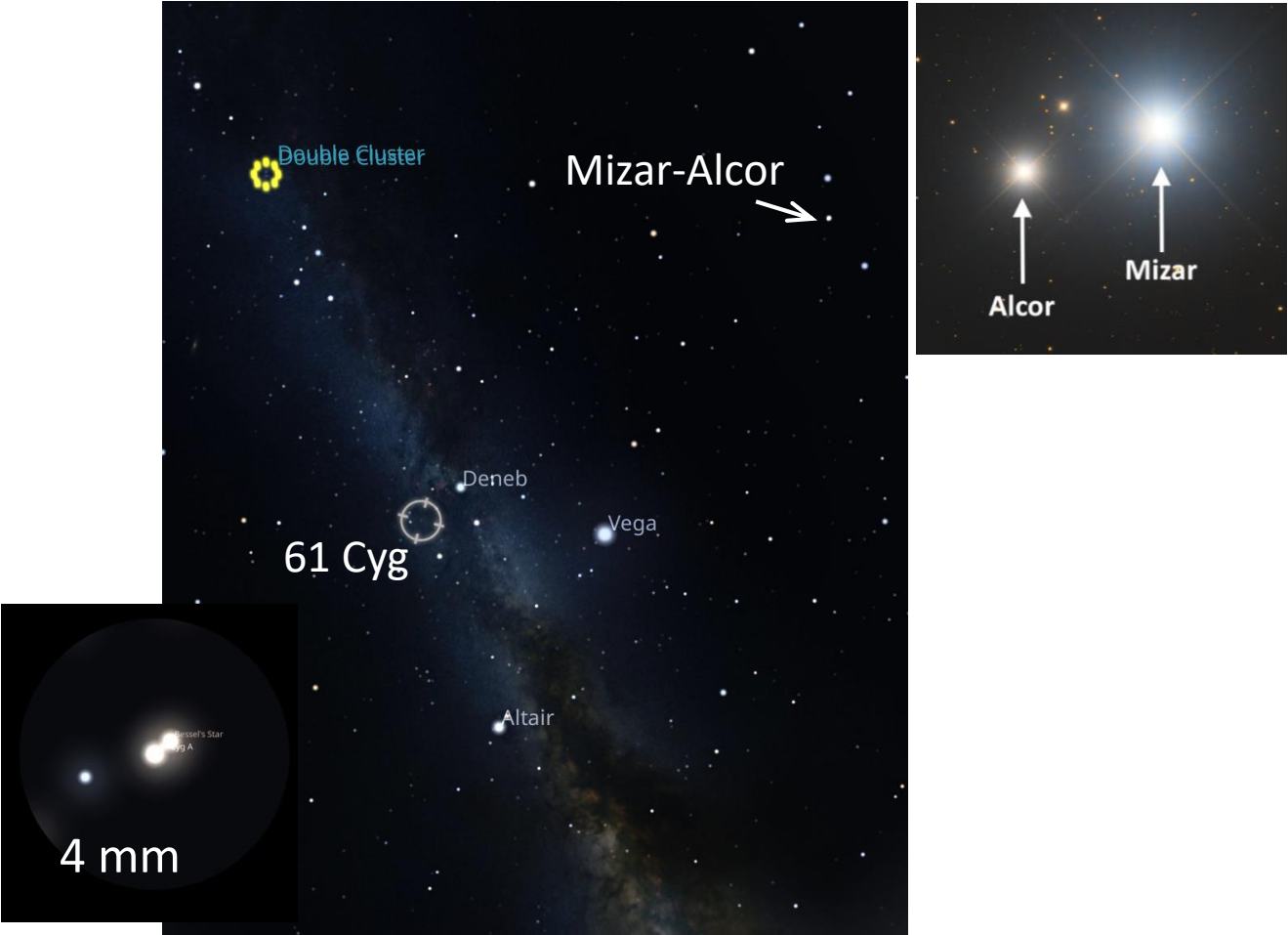
Galaxy	AR	Dec
M27	19 ^h 59 ^m 36.34 ^s	+22° 43' 16.09"
M57	18 ^h 53 ^m 35.08 ^s	+33° 01' 45.03"

Use first the 25mm to find them. M27 is already good in the 25 mm. M57 will need the 9mm (can also try the 4 mm).

Cluster	AR	Dec
Altair	19 ^h 50 ^m 47.00 ^s	+08° 52' 05.96"
Vega	18 ^h 36 ^m 56.34 ^s	+38° 47' 01.28"
Deneb	20 ^h 41 ^m 25.9 ^s	+45° 16' 49"

- What with the 114/900? M13, M15, M2, M92, M31, M81, **M27, M57**, h&χ Persei, Mizar, 61 Cygni, (M33, M51, M101)

Double Stars & Double Open Cluster



Visual Mid-High Res.

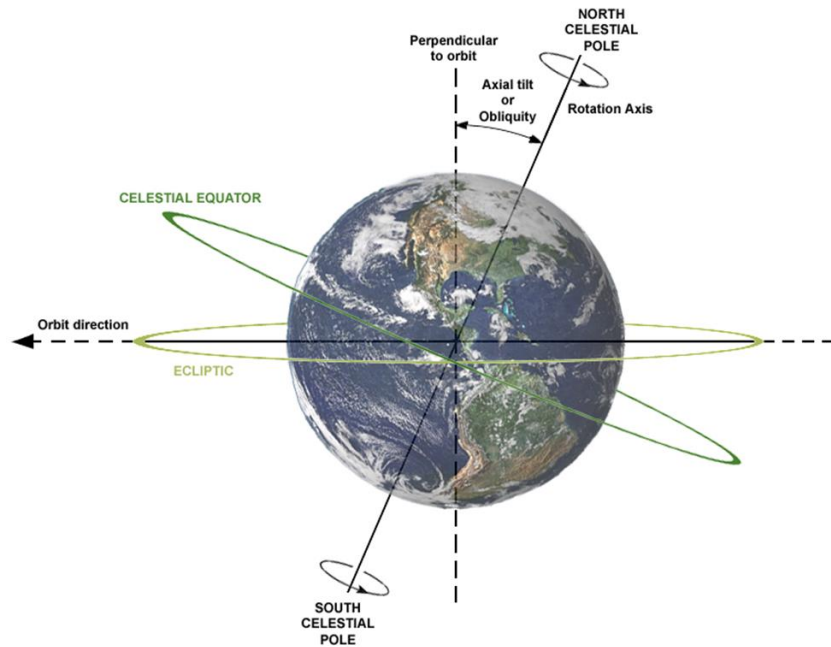
Galaxy	AR	Dec
h&χ	2 ^h 20 ^m	57° 08′
61 Cyg		

h&χ Persei and Mizar-Alcor require the 25mm. Go up to 9mm and 4mm (see insert) on 61 Cyg to see if you manage to split it...

Cluster	AR	Dec
Altair	19 ^h 50 ^m 47.00 ^s	+08° 52′ 05.96″
α Cas	00 ^h 40 ^m 30.44 ^s	+56° 32′ 14.39″
Deneb	20 ^h 41 ^m 25.9 ^s	+45° 16′ 49″

- What with the 114/900? M13, M15, M2, M92, M31, M81, M27, M57, **h&χ Persei, Mizar, 61 Cygni**, (M33, M51, M101)

Photo Skyscapes



Hours long exposure on a fixed tripod (from the Web)

- The Earth rotates so the sky apparently rotates at $\omega = 15^\circ / \text{h} = 15'' / \text{s}$ at the celestial equator.
- So, on a sensor, you get a track of $L = \omega \times F$. If $F = 50 \text{ mm}$, $L = 3.6 \mu\text{m} / \text{s}$!
 - Comparable to the typical pixel size of digital cameras.
 - It will have to move a few pixels before the elongation is noticeable.
 - An empirical rule for star fields is $T_{\text{exp}} (\text{s}) \leq 300 / F$

Photo Skyscapes

- For hand held photography $T_{exp} (s) \leq 1 / F$
 - VR can gain you a factor of 8-10
 - Still, only OK for the moon and the Sun at sunset
 - Even the Moon is small (31' diameter)
 - On the focal plane this is $\approx F / 110$ (F in mm)
 - At least $F = 300$ mm is required to see details
- It is practically impossible to go much above 1-2 s hand held, even with a wide lens (say 20 mm) and VR
- To start to obtain something you need:
 - a decent tripod and a photo head (e.g., ball head)
 - A wide fast lens ($F \leq 24$ mm, $F/\# \leq 4$)
 - A camera with good noise performance
 - Some post-processing skills and some decent software
 - For starscapes, a good dark sky
 - For the comet picture $T_{exp}=13s$ ($300/24 = 12.5$).



Single, hand held, shoot of the moon while entering into eclipse. Nikon Z6 with AF-P 70-300 @ 300mm f/5.6 ISO1000 1/50 s.



Single image from Pixhaier Teich, Harz, Germany Nikon Z 6 @ ISO 6400. 13s with 24-70 f/4 @24mm @f/4

Photo Skyscapes

- For non-tracking skyscapes you need:
 - a decent tripod and a photo head (ball heads are good)!
 - A wide fast lens ($F \leq 24 \text{ mm}$, $F/\# \leq 4$)
 - A camera with good noise performance
 - Some post-processing skills and some decent software



Composite of the Total Solar Eclipse of July 2nd 2019.
Sigma 17-50 at f/5 (19mm) on Nikon D7100
The pictures during the partial phase are taken with a
Baader astrofoil filter and 1/1000 s.
The frame during totality was obtained with a exposure
time of 1/3 s (and no filter).

Photo Skyscapes

- Without tracking, you cannot increase the exposure time but you can improve the signal to noise ratio by registering and averaging several images
 - For this picture $T_{exp}=4s$ ($300/70 = 4.3$)



Comet Neowise between light clouds
combination of 25 frames of 4s taken at ISO6400 with
a Nikon Z6 and a 24-70 f/4 @70mm @f/4

Tracking, then guiding

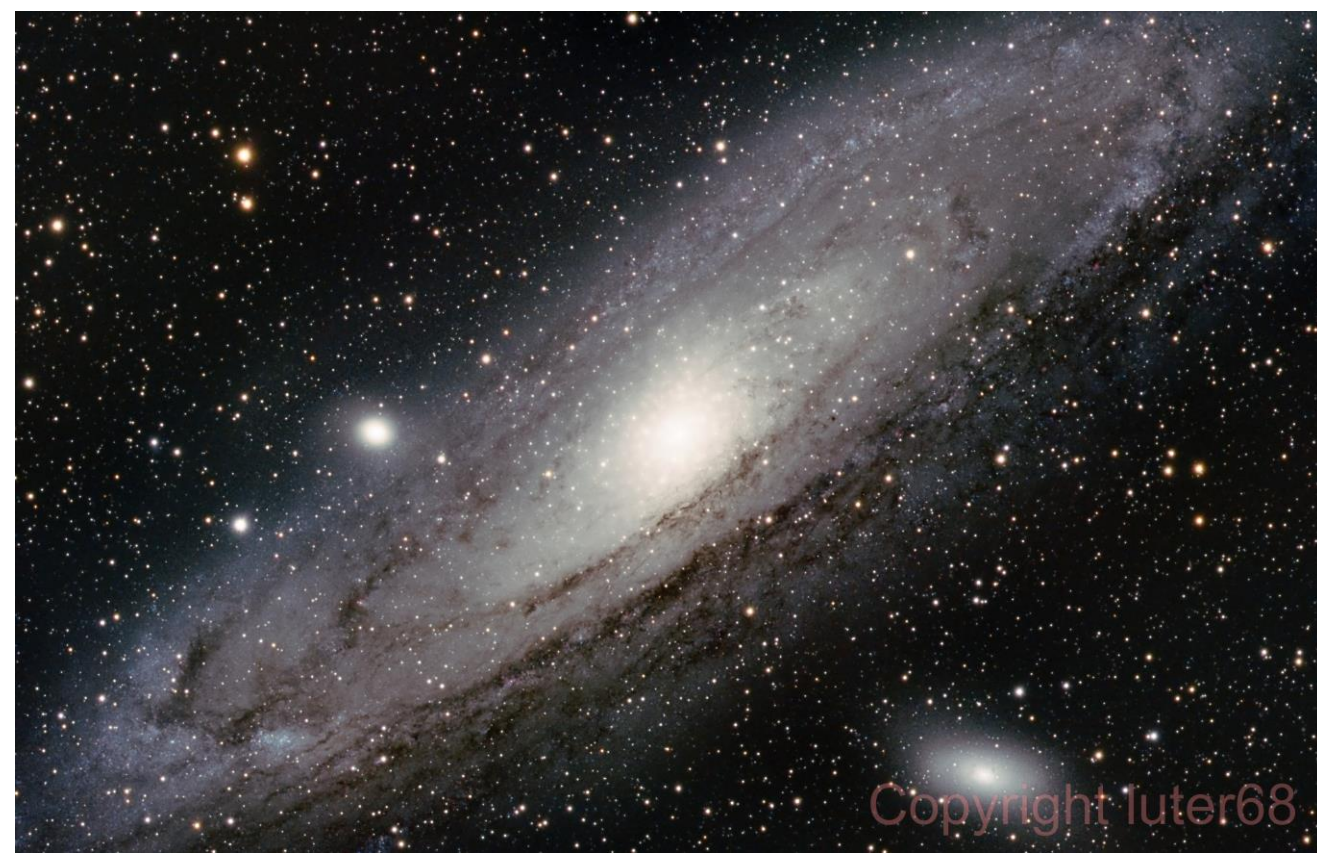
- Adding a tracking mount hugely expands possibility, depending on mount's quality
- However, cost and complexity adds up extremely fast



NGC 7000. 113 frames exposed 90 s. No guiding. Nikon 200 f/4 AI front stopped to F/5.4. Nikon D5100 with true dark hack.

Tracking, then guiding

- Guiding is the next step
 - Requires a guide camera and a guide scope
 - More complexity and cost
 - However, you can increase F and Texp
- Using a dedicated camera for astrophotography (cooled) increases a lot final quality, in particular a monochromatic camera with filters
 - Even more complexity and costs



M31. 80/400 Apo with ASI 183MM (-15 C°) on AVX

L filter: 30×180''(1h 30')

B filter: 36×90''(54')

G filter: 32×90''(48')

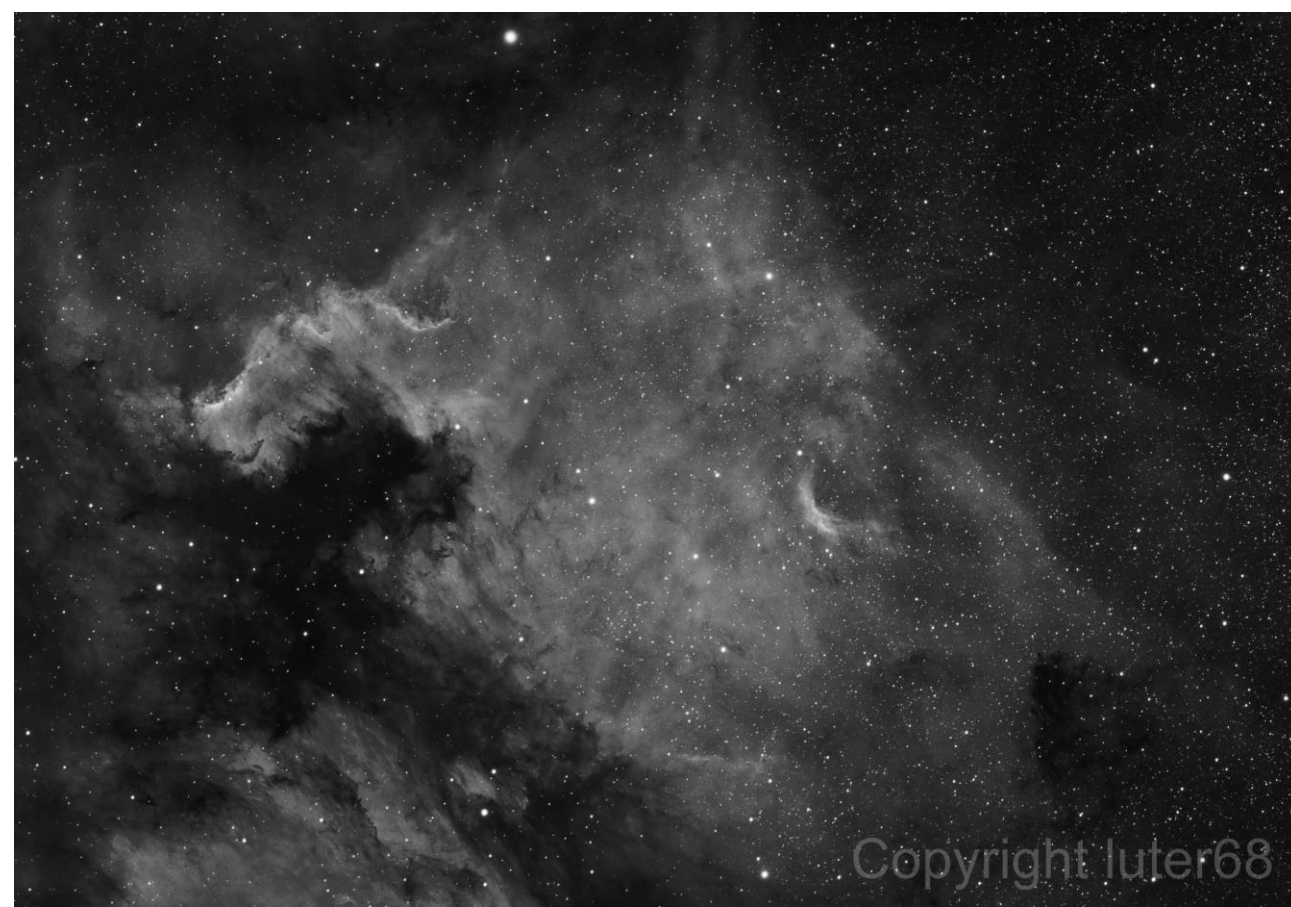
R filter: 36×90''(54')

Total integration 4h 20 min

Calibration frames: 40 darks, 25 flats, 25 dark flats (per filter)

Tracking, then guiding

- Guiding is the next step
 - Requires a guide camera and a guide scope
 - More complexity and cost
 - However, you can increase F and Texp
- Using a dedicated camera for astrophotography (cooled) increases a lot final quality, in particular a monochromatic camera with filters
 - Even more complexity and costs



NGC 7000. Nikon 200 F/4 AI front stopped at F/5.4 with ASI 183MM (-15 C°) on AVX
Ha filter: 49×600''(8h 10')
Calibration frames: 40 darks, 25 flats, 25 dark flats